

Electron-phonon scattering in topological insulator thin films

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We present a theoretical study of electron-phonon scattering effects in thin films made of a strong topological insulator. The phonons are modeled by isotropic elastic continuum theory with stress-free boundary conditions, and the interaction with the helical surface Dirac fermions is mediated by the deformation potential. We determine the temperature-dependent electrical resistivity $\rho(T)$ and the quasiparticle decay rate $\Gamma(T)$ observable in photoemission. The low- and high-temperature power laws for both quantities are obtained analytically. Detailed estimates covering the full temperature range are provided for Bi_2Se_3 .

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I. INTRODUCTION

The recently discovered state of matter called the “topological insulator” (TI) currently represents one of the most active areas in condensed matter physics.^{1,2} TIs are characterized by an insulating gap in the bulk but at the same time have gapless surface modes protected against all time-reversal-invariant (and sufficiently weak) perturbations.^{3,4} In a three-dimensional (3D) TI, these surface modes correspond to massless two-dimensional (2D) Dirac fermions, where the spin direction is in the surface plane and perpendicular to the momentum (“spin-momentum locking”). A typical reference material is Bi_2Se_3 with a bulk gap $\Delta_b \approx 0.3$ eV. The helical Dirac electron property of the TI surface state has been experimentally confirmed by spin- and angle-resolved photoemission spectroscopy (ARPES).^{1,5} Transport experiments are more difficult in that respect since the surface contribution is often masked by the residual conductivity due to impurities or defects in the bulk.^{6–8} In thin films made of TI materials, however, the bulk contribution is greatly suppressed relative to the surface contribution, rendering the latter easier to observe.

In this paper, we provide a detailed theoretical analysis of both the temperature-dependent resistivity $\rho(T)$ and the quasiparticle lifetime $\Gamma(T)$ [observable in ARPES (Refs. 9 and 10)] for a thin TI film. The approach taken here generalizes previous work for the semi-infinite geometry (with only one surface) by two of us¹¹ to the film geometry. This brings about several important changes compared to Ref. 11. In particular, the temperature dependence of both $\rho(T)$ and $\Gamma(T)$ is different in the film geometry at low temperatures, and a pronounced and interesting dependence on the film width is found. Due to the transverse quantization, the finite-width theory is more complex yet remains analytically tractable. We model the electronic part by retaining only the Dirac surface states obtained from the low-energy band structure,¹² and our theory always assumes that the Fermi level is located inside the bulk gap. We note in passing that more accurate parameter estimates are provided in Ref. 12 than in the earlier paper by the same authors,¹³ and we here adopt the new parameters in our calculations using Bi_2Se_3 as example. A similar parameter set has been published in Ref. 14. In sufficiently thin films, the hybridization of the two surface states eventually causes insulating behavior, as has recently been observed experimentally from ARPES for Bi_2Se_3 films.^{15,16} For Bi_2Se_3 , several calculations predicted^{14,17,18} a gap $\Delta(L)$ with (as a

function of the width L) oscillations superimposed on an exponential decay. Similar calculations, however, found no oscillations, with a well-established TI phase already for $L \geq 3$ quintuple layers (QLs).^{19,20} Using the parameter set of Ref. 12, we also find no evidence for oscillations in $\Delta(L)$; see Sec. II A below. For large width, one then has (upper and lower) massless Dirac fermion surfaces.²¹

Our working assumption below is that electron-phonon scattering is the dominant source of quasiparticle decay and backscattering. Electron-electron interactions are indeed expected to give only subleading corrections to the resistivity as long as $T \gtrsim 1$ mK.²² Disorder effects are more likely to compete with phonon-induced backscattering effects. However, for elevated temperatures, $T \gtrsim 100$ K, phonon effects dominate even for present-day samples, and, with higher-purity films anticipated in the future, this crossover temperature may be lowered significantly. ARPES setups allowing tests of our predictions for the quasiparticle decay rate are basically available.^{5,23–25} Other surface scattering techniques have also been applied to extract the phonon dispersions.²⁶ We here follow Ref. 11 and model the phonons using elastic continuum theory.²⁷ Since even at room temperature one effectively probes low-energy scales, we keep only long-wavelength acoustic phonon modes. For these, previous work on related materials has shown^{28,29} that isotropic elastic continuum theory provides a reasonable approximation. The phonon eigenmodes in the thin film geometry and their coupling to electronic modes have previously been determined in the context of semiconductor quantum well structures.³⁰ (Note that the semi-infinite case has been treated in Ref. 31.) We basically reproduce the phonon eigenmodes of Ref. 30, but the coupling to the helical electronic eigenstates in a TI film is different from that in the semiconductor case. Note that piezoelectric couplings are suppressed by symmetry here,²⁹ and spin-phonon-type couplings³² are also expected to be subdominant to the deformation potential taken into account below.

Most TI experiments have so far addressed only optical phonons³³ (cf. also the corresponding situation for Bi surfaces³⁴), but very recently ARPES studies reported phonon-induced broadening of the line shape in TIs.^{23–25} The observed Bi_2Se_3 electron-phonon coupling strength,²⁴ which has been extracted from the prefactor in the high-temperature quasiparticle decay rate $\Gamma \propto T$, is in good agreement with

our theoretical estimates. This indicates that the low-energy approach indeed provides a reasonable starting point. To the best of our knowledge, no detailed measurements of the temperature dependence of the TI film resistivity have been reported so far. We mention in passing that for the related case of a 2D graphene monolayer, a similar comparison of theory^{35,36} to experiment³⁷ has turned out to be successful. Remarkably, the electron-phonon coupling observed in Ref. 24 and independently estimated by us¹¹ turns out to be quite large. Under room-temperature conditions, the resulting lifetime of helical quasiparticles is therefore short, and the resistivity is rather large. This behavior is substantially different from what is found in graphene. We suspect that this is (partially) due to the different Debye temperatures in the two materials.

The structure of the remainder of this paper is as follows. In Sec. II we discuss the model for the surface states in the thin film and their coupling to the quantized phonon modes. We then turn to the calculation of the electrical resistivity in Sec. III, followed by that of the lifetime broadening in Sec. IV. The paper concludes with a brief discussion in Sec. V. Technical details of our calculations can be found in various appendixes. Note that we use units with $\hbar = 1$.

II. MODEL

In this section we describe the model employed in our study of electron-phonon scattering in a TI film. The model parameters below are chosen for Bi₂Se₃ as a concrete example. The film has infinite extension in the xy plane and the width L , where $|z| < L/2$. We start by reviewing the construction of the effective surface Hamiltonian describing the (upper and lower) electronic surface states of a TI film.

A. Electronic surface states

Keeping all terms up to second order in the momentum around the Γ point (k_x, k_y, k_z), the low-energy physics of 3D TI materials like Bi₂Se₃ or Bi₂Te₃ is well described by an effective four-band model.¹ Using the basis states $\{|P1_z^+, \uparrow\rangle, |P2_z^-, \uparrow\rangle, |P1_z^+, \downarrow\rangle, |P2_z^-, \downarrow\rangle\}$, the low-energy bulk Hamiltonian reads¹²⁻¹⁴

$$H = \begin{pmatrix} \epsilon_0 + M & -iA_1k_z & 0 & A_2k_- \\ iA_1k_z & \epsilon_0 - M & A_2k_- & 0 \\ 0 & A_2k_+ & \epsilon_0 + M & -iA_1k_z \\ A_2k_+ & 0 & iA_1k_z & \epsilon_0 - M \end{pmatrix} \quad (1)$$

with $\epsilon_0 = C + D_1k_z^2 + D_2(k_x^2 + k_y^2)$, $M = M_0 - B_1k_z^2 - B_2(k_x^2 + k_y^2)$, and $k_{\pm} = k_x \pm ik_y$. The model parameters for Bi₂Se₃ have been determined from first principles,¹²

$$\begin{aligned} M_0 &= 0.28 \text{ eV}, & C &= -0.0083 \text{ eV}, \\ A_1 &= 2.26 \text{ eV \AA}, & A_2 &= 3.33 \text{ eV \AA}, \\ B_1 &= 6.86 \text{ eV \AA}^2, & B_2 &= 44.5 \text{ eV \AA}^2, \\ D_1 &= 5.74 \text{ eV \AA}, & D_2 &= 30.4 \text{ eV \AA}^2. \end{aligned} \quad (2)$$

We may write the Hamiltonian (1) in the form $H = H_0 + H'$, where $H_0 = \begin{pmatrix} h_0(k_z) & 0 \\ 0 & h_0(k_z) \end{pmatrix}$ is the 2×2 block matrix obtained for $k_x = k_y = 0$, with

$$h_0(k_z) = \begin{pmatrix} \epsilon_0(k_z) + M_0 - B_1k_z^2 & -iA_1k_z \\ iA_1k_z & \epsilon_0(k_z) - M_0 + B_1k_z^2 \end{pmatrix}. \quad (3)$$

Note that eigenstates of H_0 have conserved spin.

In order to find the surface states in the film geometry, we follow the usual strategy¹²⁻¹⁴ and first look for general bispinor eigenstates of h_0 ,

$$h_0(k_z \rightarrow -i\partial_z)\Psi(z) = E_0\Psi(z). \quad (4)$$

The general solution to Eq. (4) reads ($j = \pm, s = \pm$)

$$\Psi(z) = \sum_{js} c_{js} e^{-s\eta_j z} \begin{pmatrix} E_0 - C + M_0 + (D_1 + B_1)\eta_j^2 \\ -sA_1\eta_j \end{pmatrix} \quad (5)$$

with arbitrary c_{js} and the inverse length scales

$$\eta_{\pm} = [(-\tilde{B} \pm \sqrt{\tilde{B}^2 - 4\tilde{A}\tilde{C}})/(2\tilde{A})]^{1/2},$$

where $\tilde{A} = D_1^2 - B_1^2$, $\tilde{B} = A_1^2 - 2[M_0B_1 + D_1(C - E_0)]$, and $\tilde{C} = (E_0 - C)^2 - M_0^2$. The Dirichlet boundary conditions defining the film geometry, $\Psi(z = \pm L/2) = 0$, then imply the transcendental equation

$$\frac{[E_0 - C + M_0 + (D_1 + B_1)\eta_+^2]\eta_-}{[E_0 - C + M_0 + (D_1 + B_1)\eta_-^2]\eta_+} = \frac{\tanh(\eta_-L/2)}{\tanh(\eta_+L/2)},$$

or the same condition with $\eta_+ \leftrightarrow \eta_-$ on the right-hand side. Numerical solution of these equations yields the Γ point energies $E_0^{(\pm)}$. The corresponding eigenstates $\Psi_{\pm}(z)$ follow from Eq. (5),

$$\Psi_{\pm}(z) = \mathcal{N}_{\pm} \begin{pmatrix} (D_1 + B_1)\Lambda_{\pm} F_{\mp}^{\pm}(z) \\ A_1 F_{\pm}^{\pm}(z) \end{pmatrix}, \quad (6)$$

where the \mathcal{N}_{\pm} are normalization constants and

$$\Lambda_{\pm} = \left[\frac{\eta_+^2 - \eta_-^2}{\eta_+ \coth^{\pm}(\eta_+L/2) - \eta_- \coth^{\pm}(\eta_-L/2)} \right]_{E_0^{(\pm)}}$$

with $\coth^+(y) = \coth(y)$ and $\coth^-(y) = \tanh(y)$. Finally, the F functions are

$$F_{\pm}^{\pm}(z) = \left[\frac{\cosh(\eta_+z)}{\cosh(\eta_+L/2)} - \frac{\cosh(\eta_-z)}{\cosh(\eta_-L/2)} \right]_{E_0^{(\pm)}},$$

where F_{\pm}^{\pm} follows with $\cosh \rightarrow \sinh$. Note that the eigenstates $\Psi_{\pm}(z)$ describe both spin directions ($\sigma = \uparrow$ and $\sigma = \downarrow$).

We now project the full Hamiltonian H to the basis spanned by the surface states (6). We define Pauli matrices $\tau_{\alpha=x,y,z}$ switching between the two solutions $\Psi_{\tau=\pm}(z)$ and Pauli matrices σ_{α} in spin space, and use τ_0 and σ_0 as identities. With the energy scales

$$E_0 = \frac{E_0^{(+)} + E_0^{(-)}}{2}, \quad \Delta = E_0^{(+)} - E_0^{(-)}, \quad (7)$$

the low-energy (“surface”) Hamiltonian resulting from this projection reads

$$H_{\text{eff}} = E_0 \tau_0 \sigma_0 + \frac{\Delta}{2} \tau_z \sigma_0 - A_2 W \tau_x (k_x \sigma_x + k_y \sigma_y) + O(\mathbf{k}^2), \quad (8)$$

where $W = \langle \Psi_+ | \Psi_- \rangle$. The parameter $\Delta(L)$ is precisely the surface state gap described in the Introduction. For the parameters (2), η_- is always real. However, η_+ is real for large L but purely imaginary for small L . In any case, we find that W is always real and positive.

Since H_{eff} commutes with $\tau_z \sigma_z$, it can readily be diagonalized by the unitary transformation $U(\mathbf{k}) = \text{diag}(U_+, U_-)$, where $\mathbf{k} = (k_x, k_y)$ and the $U_{v=\pm}(\mathbf{k})$ are 2×2 matrices in spin space, with v denoting the eigenvalue of $\tau_z \sigma_z$. With $\tan \alpha = 2A_2 W |\mathbf{k}| / \Delta$ and $\tan \theta = k_y / k_x$, we find

$$U_{v=+} = \begin{pmatrix} e^{-i\theta/2} \cos(\alpha/2) & e^{-i\theta/2} \sin(\alpha/2) \\ -e^{i\theta/2} \sin(\alpha/2) & e^{i\theta/2} \cos(\alpha/2) \end{pmatrix}, \quad (9)$$

$$U_{v=-} = \begin{pmatrix} -e^{-i\theta/2} \sin(\alpha/2) & e^{-i\theta/2} \cos(\alpha/2) \\ e^{i\theta/2} \cos(\alpha/2) & e^{i\theta/2} \sin(\alpha/2) \end{pmatrix}.$$

Switching to second-quantized notation, the eigenstates of H_{eff} correspond to helical fermions with annihilation operator

$$c_{\mathbf{k}, v, s} = \sum_{\sigma} [U_v(\mathbf{k})]_{\sigma s}^* d_{\mathbf{k}, \tau=v\sigma, \sigma}, \quad (10)$$

where $d_{\mathbf{k}, \tau\sigma}$ annihilates a spin- σ electron with in-plane momentum \mathbf{k} in the transverse state $\Psi_{\tau}(z)$. The low-energy electronic Hamiltonian (including the chemical potential μ) then takes the final form

$$H_{\text{el}} = \sum_{\mathbf{k}; v, s=\pm} \epsilon_{\mathbf{k}, s} c_{\mathbf{k}, v, s}^{\dagger} c_{\mathbf{k}, v, s}, \quad (11)$$

where the dispersion relation is

$$\epsilon_{\mathbf{k}, \pm} = E_0 - E_0^{\infty} - \mu \pm \frac{\Delta}{2} \sqrt{1 + (2A_2 W / \Delta)^2 \mathbf{k}^2}. \quad (12)$$

We here choose the zero of energy by setting $E_0^{\infty} = C + D_1 M_0 / B_1 = \lim_{L \rightarrow \infty} E_0^{(\pm)}$. For the parameters (2), we find $E_0^{\infty} \simeq 0.22$ eV. Moreover, for $L \rightarrow \infty$, the length scales η_{\pm}^{-1} are given by $\eta_+^{-1} \simeq 12.3$ Å and $\eta_-^{-1} \simeq 1.9$ Å. For $kL \gg 1$, the dispersion relation (12) is linear, with Fermi velocity $v_F \simeq 2.77 \times 10^5$ m/s. Note that the index $s = \pm$ in Eq. (11) does not correspond to spin anymore.

Similarly, the particle density operator $\hat{n}(\mathbf{r}, z)$ with $\mathbf{r} = (x, y)$ is written in terms of the $d_{\mathbf{k}, \tau\sigma}$ operators,

$$\hat{n}(\mathbf{r}, z) = \sum_{\mathbf{k}, \tau, \sigma} e^{-i\mathbf{q}\cdot\mathbf{r}} \rho_{\tau}(z) d_{\mathbf{k}+\mathbf{q}, \tau\sigma}^{\dagger} d_{\mathbf{k}, \tau\sigma}. \quad (13)$$

Using Eq. (10), the density operator (13) can be transformed to the helical basis. We show the single-particle densities for the surface states [Eq. (6)],

$$\rho_{\tau}(z) = [\Psi_{\tau}^{\dagger} \cdot \Psi_{\tau}](z), \quad (14)$$

in Fig. 1 for a film thickness of $L = 4$ QL, where 1 QL $\simeq 9.5$ Å for Bi_2Se_3 .³⁸ This demonstrates that already for quite thin films Eq. (6) describes surface states. Note that $\rho_{\tau}(z)$ is an even function of z . The inset of Fig. 1 shows the numerically

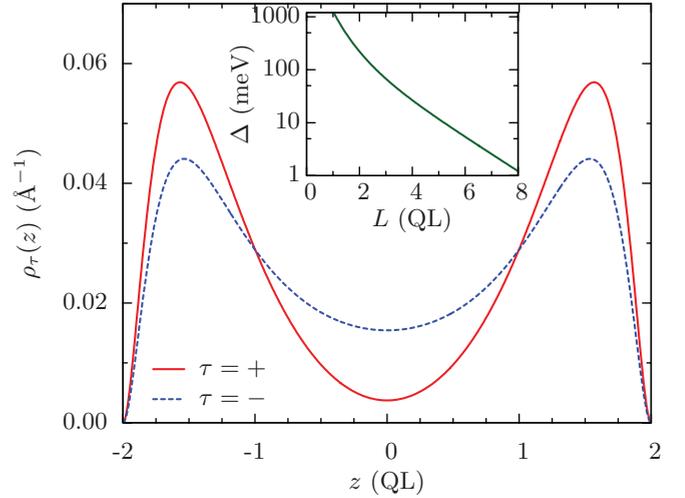


FIG. 1. (Color online) Electronic eigenstates for Bi_2Se_3 from Eqs. (1) and (2). Main panel: Densities $\rho_{\tau}(z)$ in Eq. (14) for $L = 4$ QL. Inset: Gap Δ vs thickness L . Note the semilogarithmic scale.

obtained gap $\Delta(L)$, demonstrating the absence of oscillatory behavior for the parameters (2) as well as the exponential decay of $\Delta(L)$ due to the exponentially vanishing overlap of both surface states. We note in passing that, for the parameters in Ref. 13, Eq. (7) instead predicts an oscillatory decay of $\Delta(L)$.

B. Phonon model

We now discuss the long-wavelength acoustic phonon modes in the TI film. We employ isotropic elastic continuum theory, where the longitudinal (c_l) and transverse (c_t) sound velocities correspond to the two Lamé constants.²⁷ In Bi_2Se_3 , they are given^{39,40} by $c_l \simeq 2900$ m/s and $c_t \simeq 1700$ m/s, respectively. Moreover, the mass density is $\rho_M = 7680$ kg/m³.⁴¹ In order to model the film geometry, we impose stress-free boundary conditions²⁷ at $z = \pm L/2$. The quantized phonon eigenmodes for this problem have been determined in Ref. 30. For convenience, we briefly summarize the results next.

Different phonon modes are labeled by a set of quantum numbers $\Lambda = (\mathbf{q}, \lambda, n)$, where $\mathbf{q} = (q_x, q_y)$ is the surface momentum, $\lambda \in (H, S, A)$ denotes the mode type, and $n \in \mathbb{N}$ is a branch index corresponding to the quantization of transverse momentum. The horizontal shear mode ($\lambda = H$) decouples from all other modes and does not generate a deformation potential,³⁰ and we do not discuss this mode further. We are left with transversally symmetric (dilatational, $\lambda = S$) and antisymmetric (flexural, $\lambda = A$) phonons. Denoting the dispersion relation of a given phonon mode Λ by Ω_{Λ} (see below) and the surface area by \mathcal{A} , the displacement field operator is

$$\mathbf{U}(\mathbf{r}, z, t) = \sum_{\Lambda} \frac{e^{i(\mathbf{q}\cdot\mathbf{r} - \Omega_{\Lambda} t)}}{\sqrt{2\rho_M \mathcal{A} \Omega_{\Lambda}}} \mathbf{u}_{\Lambda}(z) b_{\Lambda} + \text{H.c.}, \quad (15)$$

where b_{Λ} is a bosonic annihilation operator and the noninteracting phonon Hamiltonian is

$$H_{\text{ph}} = \sum_{\Lambda} \Omega_{\Lambda} (b_{\Lambda}^{\dagger} b_{\Lambda} + 1/2). \quad (16)$$

The orthonormal eigenmodes $\mathbf{u}_\Lambda(z)$ in Eq. (15) describe linear combinations of $e^{\pm ik_{l,t}z}$ waves, where

$$k_{l,t} = \sqrt{(\Omega_\Lambda/c_{l,t})^2 - q^2}; \quad (17)$$

$k_{l,t} = i\kappa_{l,t}$ with $\kappa_{l,t} = \sqrt{q^2 - (\Omega_\Lambda/c_{l,t})^2}$ for $\Omega_\Lambda < c_{l,t}q$. Writing $\mathbf{u}_\Lambda(z)$ in the form

$$\mathbf{u}(z) = \left(iq\phi_l - \frac{d\phi_l}{dz} \right) \hat{e}_q + \left(\frac{d\phi_l}{dz} + iq\phi_l \right) \hat{e}_z, \quad (18)$$

where $\hat{e}_q = \mathbf{q}/q$ and

$$\phi_{l,t} = a_{l,t} \cos(k_{l,t}z) + b_{l,t} \sin(k_{l,t}z), \quad (19)$$

the stress-free boundary conditions at $z = \pm L/2$ yield

$$\begin{aligned} 2iq \frac{d\phi_l}{dz} - (q^2 - k_t^2)\phi_l &= 0, \\ 2iq \frac{d\phi_t}{dz} + (q^2 - k_l^2)\phi_t &= 0. \end{aligned} \quad (20)$$

Since both equations have to be fulfilled at $z = \pm L/2$, we have four linear equations for the four unknown parameters $(a_{l,t}, b_{l,t})$. Setting the corresponding determinant to zero, we obtain the following two possibilities. First, for symmetric modes ($\lambda = S$), we have the condition

$$\begin{aligned} (q^2 - k_t^2)^2 \cos(k_l L/2) \sin(k_t L/2) \\ + 4q^2 k_l k_t \sin(k_l L/2) \cos(k_t L/2) &= 0. \end{aligned} \quad (21)$$

Numerical solution of this transcendental equation gives the quantized set of dilatational phonon frequencies $\Omega_{\Lambda=(q,S,n)}$. The corresponding eigenvector $\mathbf{u}_\Lambda(z)$ follows from Eqs. (18) and (19) with $a_t = b_l = 0$ and

$$a_l = \frac{2\mathcal{N}_S q}{\cos(k_l L/2)}, \quad b_t = \frac{i\mathcal{N}_S (q^2 - k_t^2)}{k_t \cos(k_t L/2)}. \quad (22)$$

Second, for antisymmetric modes ($\lambda = A$), we arrive again at the condition in Eq. (21) but with the exchange $\cos \leftrightarrow \sin$. Solving that equation yields the set $\Omega_{\Lambda=(q,A,n)}$ of quantized flexural phonon modes. The eigenvector $\mathbf{u}_\Lambda(z)$ follows again from Eqs. (18) and (19), where now $a_l = b_t = 0$ and

$$b_l = \frac{2\mathcal{N}_A q}{\sin(k_l L/2)}, \quad a_t = \frac{-i\mathcal{N}_A (q^2 - k_t^2)}{k_t \sin(k_t L/2)}. \quad (23)$$

The normalization factors $\mathcal{N}_{\lambda=S,A}$ appearing in Eqs. (22) and (23) are given in Appendix A.

Numerical solution of Eq. (21) yields the spectrum Ω_Λ for the symmetric mode ($\lambda = S$). The result is shown in Fig. 2. We distinguish three different regions, namely, a case where both k_l and k_t are purely imaginary (region I), a case where only k_l is purely imaginary but k_t is real (region II), and finally a case where both k_l and k_t are real (region III). We observe from Fig. 2 that the $n = 1$ mode is the finite-width analog of the well-known Rayleigh surface mode.^{27,31} For the semi-infinite geometry, the Rayleigh mode is the lowest-lying phonon.⁴² It has the dispersion relation

$$\Omega = c_R q, \quad c_R \simeq 0.92 c_t. \quad (24)$$

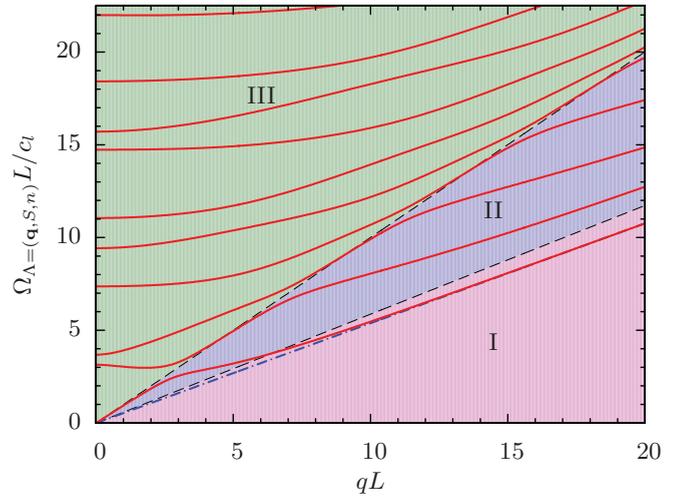


FIG. 2. (Color online) Phonon dispersion relation Ω_Λ vs q for the symmetric ($\lambda = S$) mode (red solid curves). Shown are the ten lowest branches corresponding to the index n . Dashed lines separate regions I, II, and III (see text). The dash-dotted line gives the dispersion relation in Eq. (24); note that the $n = 1$ mode coincides with the Rayleigh mode for $qL \gg 1$.

In fact, for $qL \gg 1$, both Eq. (21) and the corresponding equation for $\lambda = A$ reduce to

$$(q^2 + \kappa_t^2)^2 = 4q^2 \kappa_t \kappa_l.$$

As discussed in Ref. 27, this equation readily yields the sound velocity c_R of the Rayleigh mode.

C. Electron-phonon coupling

The dominant coupling of the above phonon modes to the electronic surface states comes from the deformation potential,¹¹ which couples the local electronic density $\hat{n}(\mathbf{r}, z)$ [Eq. (13)] to the divergence of the displacement vector, $\nabla \cdot \mathbf{U}(\mathbf{r}, z)$; see Eq. (15). Since the surface state density $\rho_\tau(z)$ in Eq. (14) is even in z , the antisymmetric phonon mode ($\lambda = A$) does not couple to the surface states. We therefore keep only the symmetric phonon mode from now on (and omit the index $\lambda = S$). Transforming Eq. (13) to the helical basis [see Eq. (10)], the second-quantized electron-phonon coupling Hamiltonian reads

$$H_{e\text{-ph}} = \frac{\alpha}{\sqrt{\mathcal{A}}} \sum_{\mathbf{q}, \mathbf{k}, n; v, s, s'} M_{\mathbf{k}, \mathbf{q}, n}^{(v, s, s')} b_{\mathbf{q}, n} c_{\mathbf{k}+\mathbf{q}, vs}^\dagger c_{\mathbf{k}, vs'} + \text{H.c.}, \quad (25)$$

where the M matrix elements involve the unitary matrices $[U_v(\mathbf{k})]_{\sigma\sigma'}$ in Eq. (9),

$$\begin{aligned} M_{\mathbf{k}, \mathbf{q}, n}^{(v, s, s')} &= -\frac{1}{\sqrt{2\rho_M \Omega_{q,n}}} \left(\frac{\Omega_{q,n}}{c_l} \right)^2 \\ &\times \sum_{\sigma} [U_v(\mathbf{k} + \mathbf{q})]_{\sigma\sigma'}^* [U_v(\mathbf{k})]_{\sigma\sigma'} \\ &\times \int_{-L/2}^{L/2} dz \rho_{\tau=vs}(z) \phi_l(z), \end{aligned} \quad (26)$$

with the phonon dispersion $\Omega_{q,n}$ in Fig. 2; ϕ_l is given by Eqs. (19) and (22). The deformation potential strength α in Eq. (25) can be estimated as follows. The high-temperature behavior of the on-shell imaginary part of the electronic self-energy is (see Sec. IV)

$$\text{Im } \Sigma(k, T) = -\pi \lambda_k k_B T, \quad (27)$$

which allows experimental extraction of the dimensionless effective electron-phonon coupling constant λ_k . The relation (27) has been observed for Bi₂Se₃ in ARPES experiments,²⁴ and $\lambda = 0.25 \pm 0.05$ has been measured. In these experiments, the Fermi level was near the bottom of the conduction band, $\mu \simeq 0.28$ eV, and k in Eq. (27) corresponds to energies ≈ 50 to 100 meV above the Dirac point. With λ_k computed within our model (see Sec. IV), the observed value for λ corresponds to $\alpha = (30 \pm 8)$ eV. We employ the value $\alpha = 30$ eV below.

The total Hamiltonian employed in the following sections is then given by $H = H_{\text{el}} + H_{\text{ph}} + H_{e\text{-ph}}$; see Eqs. (11), (16), and (25). We first address the phonon-induced resistivity ρ in Sec. III and then turn to the quasiparticle lifetime in Sec. IV.

III. RESISTIVITY

Here we discuss the T -dependent phonon contribution to the electrical resistivity ρ in the TI film, using the Hamiltonian described in Sec. II. As explained in Sec. II C, only symmetric (dilatational) phonon modes can cause a finite resistivity for the low-energy surface states within the bulk gap. We compute ρ within the framework of the linearized Boltzmann equation,⁴³ which has also been employed previously for the related graphene case.^{35,36} The resulting quasiclassical estimate for ρ is valid³⁵ as long as ρ is small compared to the resistance quantum, $\rho \ll h/e^2 \simeq 25.8$ k Ω . We sketch the standard derivation^{30,35,36,44} for ρ in Appendix B. The result takes the form

$$\frac{1}{\rho} = \frac{e^2}{2} \sum_{\nu, s=\pm} \int \frac{d\mathbf{k}}{(2\pi)^2} v_{\mathbf{k},s}^2 \tau_\nu(\epsilon_{\mathbf{k},s}) [-\partial_\epsilon n_F(\epsilon_{\mathbf{k},s})], \quad (28)$$

where the dispersion relation for helical fermions [Eq. (12)] defines the group velocity $v_{\mathbf{k},s} = \hat{e}_k \cdot \nabla_{\mathbf{k}} \epsilon_{\mathbf{k},s}$. Moreover, $n_F(\epsilon)$ is the Fermi function, and the energy-dependent electron-phonon transport scattering rate (inverse time) is

$$\frac{1}{\tau_\nu(\epsilon_{\mathbf{k},s})} = \sum_{\mathbf{q}, s'} \left(1 - \frac{v_{\mathbf{k}+\mathbf{q},s'}}{v_{\mathbf{k},s}} \cos \theta_{\mathbf{k},\mathbf{q}} \right) \times \frac{1 - n_F(\epsilon_{\mathbf{k}+\mathbf{q},s'})}{1 - n_F(\epsilon_{\mathbf{k},s})} W_{(\mathbf{k},\nu s) \rightarrow (\mathbf{k}+\mathbf{q},\nu s')}, \quad (29)$$

where $\theta_{\mathbf{k},\mathbf{q}}$ is the angle between \mathbf{k} and $\mathbf{k} + \mathbf{q}$, and the transition probabilities are obtained from Fermi's golden rule. Using Eq. (25), we find

$$W_{(\mathbf{k},\nu s) \rightarrow (\mathbf{k}+\mathbf{q},\nu s')} = \frac{2\pi\alpha^2}{\mathcal{A}} \sum_{n,\nu=\pm} \nu n_B(\nu\Omega_{q,n}) |M_{\mathbf{k},\mathbf{q},n}^{(s',s)}|^2 \times \delta(\epsilon_{\mathbf{k},s} + \nu\Omega_{q,n} - \epsilon_{\mathbf{k}+\mathbf{q},s'}), \quad (30)$$

where $n_B(\epsilon)$ is the Bose function. While the M matrix elements (26) depend on the index $\nu = \pm$, we note that $|M|^2$ and therefore the transition probabilities W are ν independent. This also implies that τ_ν does actually not depend on ν .

With the polar angle θ between \mathbf{k} and \mathbf{q} , such that

$$\cos \theta_{\mathbf{k},\mathbf{q}} = \frac{k + q \cos \theta}{\sqrt{k^2 + q^2 + 2kq \cos \theta}}, \quad (31)$$

the angular integration in Eq. (29) can be encapsulated in the ‘‘transport Eliashberg function’’ (see also Ref. 11)

$$\mathcal{F}_{k,n,s}^{(\nu)}(q) = \sum_{s'} \int_{-\pi}^{\pi} \frac{d\theta}{2\pi} \left[1 - \frac{v_{\mathbf{k}+\mathbf{q},s'}}{v_{\mathbf{k},s}} \cos \theta_{\mathbf{k},\mathbf{q}} \right] |M_{\mathbf{k},\mathbf{q},n}^{(s',s)}|^2 \times \delta(\epsilon_{\mathbf{k},s} + \nu\Omega_{q,n} - \epsilon_{\mathbf{k}+\mathbf{q},s'}). \quad (32)$$

This allows us to write the momentum relaxation rate (29) in the form

$$\frac{1}{\tau(\epsilon_{\mathbf{k},s})} = \alpha^2 \sum_{n,\nu} \int_0^\infty q dq \mathcal{F}_{k,n,s}^{(\nu)}(q) \nu n_B(\nu\Omega_{q,n}) \times \frac{1 - n_F(\epsilon_{\mathbf{k},s} + \nu\Omega_{q,n})}{1 - n_F(\epsilon_{\mathbf{k},s})}. \quad (33)$$

The θ integration in Eq. (32) can then be carried out analytically. We quote the (lengthy) result, which is useful when computing \mathcal{F} numerically, in Appendix C. For low temperatures, the quasielastic approximation $\Omega_{q,n} \ll \sqrt{(\Delta/2)^2 + (A_2 W k)^2}$ is applicable and allows us to simplify the full result for \mathcal{F} to the ν -independent form

$$\mathcal{F}_{k,n,s}(q) = \Theta(2k - q) \frac{1}{\pi \sqrt{(2k/q)^2 - 1}} \times \frac{\sqrt{(\Delta/2)^2 + (A_2 W k)^2}}{(A_2 W k)^2} |M_{\mathbf{k},\mathbf{q},n}^{(s,s)}|^2 \Big|_{\theta_0}, \quad (34)$$

where $\theta = \theta_0$ (see Appendix C) determines the polar angle between \mathbf{k} and \mathbf{q} appearing in the matrix element M , and the Heaviside function is denoted by $\Theta(y)$. Note that there is no contribution from interband transitions at low temperatures.

The crossover temperature from the low- to the high-temperature behavior in this system is set^{11,35} by the Bloch-Grüneisen temperature

$$T_{\text{BG}} = 2k_F c_R / k_B, \quad (35)$$

with the Rayleigh velocity c_R in Eq. (24). $k_F(L)$ is defined by $\epsilon_{k_F, s=\pm} = 0$ with the dispersion relation (12). For $T \ll T_{\text{BG}}$, the \mathcal{F} function can be approximated by the quasielastic expression [Eq. (34)]. It receives the dominant contribution from the $n = 1$ branch corresponding to the Rayleigh surface phonon. For small q , we find $\Omega_{q,n=1} = c_s q$ with $c_s = 2754$ m/s (which is slightly below c_l); see also Fig. 2. In addition, we have $\phi_l(z) = 2(c_l/c_s)^2 / (q\sqrt{L})$ and $c_s \ll \min(|v_{k_F}|, A_2 W)$, leading to

$$\mathcal{F}_{k_F,1,\pm}(q) = \frac{(c_l/c_s)^4}{\pi \rho_M |v_{k_F}| c_s k_F^2} \frac{q^2}{L}.$$

This allows us to perform all remaining integrations and yields a T^4 law for the resistivity at low temperatures,

$$\rho(T \ll T_{\text{BG}}) = \frac{h}{e^2} A \left(\frac{T}{T_{\text{BG}}} \right)^4, \quad (36)$$

where the dimensionless prefactor A is

$$A = \frac{8\gamma k_F \alpha^2}{\pi \rho_M v_{k_F}^2 c_s} \left(\frac{c_t c_R}{c_l c_s} \right)^4 \frac{1}{L}, \quad (37)$$

$$\gamma = \left[\int_{-\infty}^{\infty} dx \frac{2e^x}{[(\pi^2 + x^2)(e^x + 1)]^2} \right]^{-1} \simeq 68.4295.$$

For $L \rightarrow \infty$, A obviously vanishes. This suggests that for elevated temperatures (but still $T < T_{\text{BG}}$) and finite L , the T^4 law is replaced by the L -independent $\rho \propto T^5$ law found in Ref. 11. We can estimate the crossover temperature T_c as follows. For $T < T_{\text{BG}}$, we expect an expansion of the form

$$(e^2/h)\rho = A(T/T_{\text{BG}})^4 + \frac{B}{4}(T/T_{\text{BG}})^5,$$

with $A \propto 1/L$ in Eq. (36) and the L -independent constant B given in Ref. 11. The crossover from the T^4 law (for $T \lesssim T_c$) to the T^5 law (for $T_c \lesssim T < T_{\text{BG}}$) thus happens around the temperature $T_c = (4A/B)T_{\text{BG}}$. This gives $T_c \simeq 0.14T_{\text{BG}}/(k_F L)$, which is independent of the chemical potential since $T_{\text{BG}} \propto k_F$. For $L = 4$ QL, we obtain $T_c \approx 0.9$ K. The T^4 law can thus only be observed for very thin and clean TI films.

In the opposite high-temperature limit, essentially all phonon branches indexed by n contribute to the transport Eliashberg function (32); see Appendix C. Then the relaxation rate $\tau^{-1}(\epsilon_{\mathbf{k},s})$ in Eq. (33) is basically a linear function of the energy. Since the linear term does not contribute to ρ after integration in Eq. (28), we obtain the approximation $1/\rho \simeq (e^2/h)v_{k_F}k_F\tau(\epsilon = 0)$, where Eq. (33) yields the linear high-temperature law

$$\rho(T \gg T_{\text{BG}}) = \frac{h}{e^2} C \frac{T}{T_{\text{BG}}} \quad (38)$$

with the dimensionless prefactor

$$C = \frac{2\alpha^2 c_R}{v_{k_F}} \sum_{n,v=\pm} \int_0^{\infty} q dq \frac{\mathcal{F}_{k_F,n,+}^{(v)}(q)}{\Omega_{q,n}}. \quad (39)$$

Next we show the full temperature dependence of ρ obtained numerically for a fixed width $L = 4$ QL and several values of the chemical potential μ ; see Fig. 3. In that case, when measured relative to E_0^∞ , we have $E_0^+ \simeq 16$ meV and $\Delta/2 \simeq 13$ meV. For the lowest μ in Fig. 3, the Fermi level is thus located inside the surface gap and one has a very large resistivity, where the quasiclassical approach is not reliable in any case. For low temperatures $T < T_{\text{BG}}$, the analytical result (36) with $\rho \propto T^4$ is nicely reproduced by numerics. In this temperature regime, only the Rayleigh mode ($n = 1$) is relevant, similar to what one finds in the semi-infinite geometry.¹¹ In the high-temperature limit, both the $\rho \propto T$ scaling and the prefactor C in Eq. (39) are also consistent with our numerical findings.

Finally, Fig. 4 shows the width (L) dependence of ρ at fixed chemical potential and for several T . Two noteworthy observations can be drawn from Fig. 4: First, for low temperatures we observe a ‘‘dip’’ in Fig. 4, where $\rho(L) < \rho(L \rightarrow \infty)$ for intermediate values of L . Second, for $L \rightarrow \infty$, $\rho(L)$ approaches 1/4 of the single-surface value $\rho_\infty(T)$ obtained for the semi-infinite geometry.¹¹ Naively, we would expect $\rho(L \rightarrow \infty) = \rho_\infty/2$ because of the presence of two surfaces in

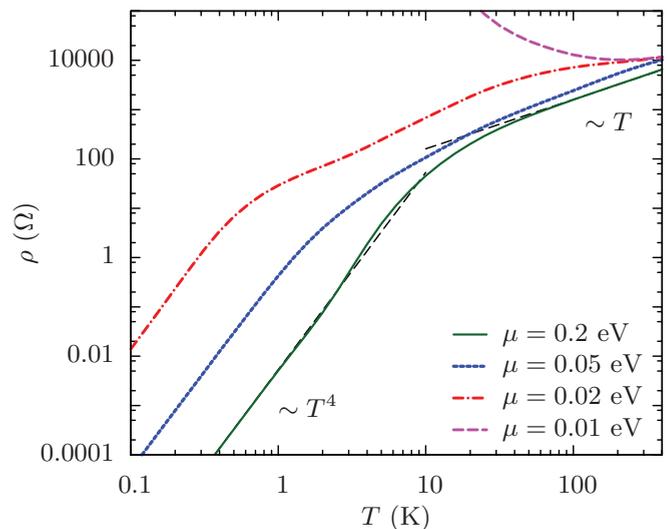


FIG. 3. (Color online) Phonon contribution to the resistivity ρ vs temperature T for a TI film of width $L = 4$ QL and several values of the chemical potential μ . Dashed straight lines indicate the analytical results for low [Eq. (36)] and high [Eq. (38)] temperatures. Note the double-logarithmic scale.

the film geometry. This discrepancy indicates that the $L \rightarrow \infty$ limit is singular, and it is not possible to really decouple both surfaces in such an interacting system; see also Ref. 21 for a related discussion.

IV. LIFETIME BROADENING

Next we discuss the quasiparticle lifetime (inverse decay rate) for the surface fermions in the TI film due to their coupling to phonons [see $H_{e\text{-ph}}$ in Eq. (25)], which implies a finite linewidth of ARPES spectral features. The decay rate $\Gamma_k(T) = -2 \text{Im} \Sigma$ follows from the imaginary part of the on-shell

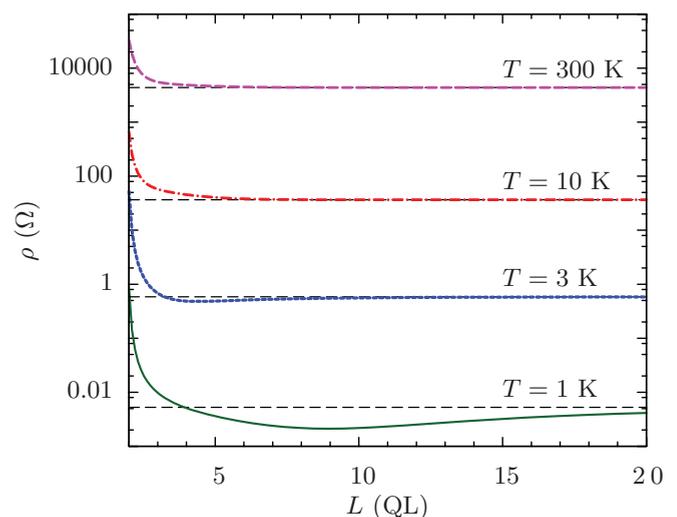


FIG. 4. (Color online) Width (L) dependence of the phonon contribution to the resistivity ρ for $\mu = 0.2$ eV and several temperatures. The dashed horizontal line indicates one-quarter of the resistivity $\rho_\infty(T)$ in the semi-infinite geometry with otherwise identical parameters (Ref. 11).

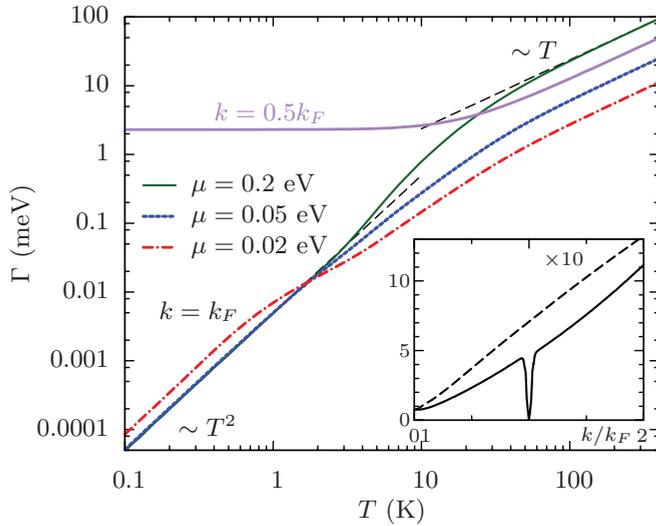


FIG. 5. (Color online) Main panel: T dependence of the decay rate Γ of a TI film of width $L = 4$ QL for $k = k_F$ and $k = 0.5k_F$. For $k = 0.5k_F$, only the $\mu = 0.2$ eV result is displayed. Dashed lines indicate the low- and high-temperature laws ($\Gamma \propto T^2$ and $\Gamma \propto T$), respectively. Inset: k dependence of Γ for $\mu = 0.2$ eV and two different temperatures: $T = 3$ K (solid line) and $T = 300$ K (dashed line; the shown result has to be multiplied by 10).

self-energy $\Sigma_{s=+}(\mathbf{k}, \omega = \epsilon_{\mathbf{k}, s=+})$ (see Fig. 5).⁴⁵ Expanding up to second order in $H_{e\text{-ph}}$, the “rainbow” diagram yields the self-energy

$$\Sigma_s(\mathbf{k}, \omega) = \alpha^2 \sum_{n, s'} \int \frac{d\mathbf{q}}{(2\pi)^2} |M_{\mathbf{k}, \mathbf{q}, n}^{(s', s)}|^2 \times \sum_{\nu = \pm} v \frac{n_B(\nu \Omega_{q, n}) + n_F(\epsilon_{\mathbf{k} + \mathbf{q}, s'})}{\omega + i0^+ + \nu \Omega_{q, n} - \epsilon_{\mathbf{k} + \mathbf{q}, s'}}. \quad (40)$$

Introducing the Eliashberg function $F_{k, n, s}^{(\nu)}(q)$ exactly as the transport Eliashberg function \mathcal{F} in Eq. (32) but without the factor $[1 - (v_{\mathbf{k} + \mathbf{q}, s'} / v_{\mathbf{k}, s}) \cos \theta_{\mathbf{k}, \mathbf{q}}]$, the quasiparticle decay rate follows as

$$\Gamma_k(T) = \alpha^2 \sum_{n, \nu} \int_0^\infty q dq F_{k, n, +}^{(\nu)}(q) \times [n_B(\Omega_{q, n}) + n_F(\Omega_{q, n} + \nu \epsilon_{k, +})]. \quad (41)$$

Expanding this result for high temperatures $T \gg T_{\text{BG}}$, as in Sec. III, yields [see also Eq. (27)] a linear T dependence,

$$\Gamma_k(T \gg T_{\text{BG}}) = 2\pi \lambda_k k_B T, \quad (42)$$

$$\lambda_k = \frac{\alpha^2}{2\pi} \sum_{n, \nu} \int_0^\infty q dq \frac{F_{k, n, +}^{(\nu)}(q)}{\Omega_{q, n}}.$$

The L dependence of λ_k is shown for $k = k_F$ in Fig. 6. We observe an oscillatory dependence, with a saturation at one-half of the corresponding semi-infinite result.

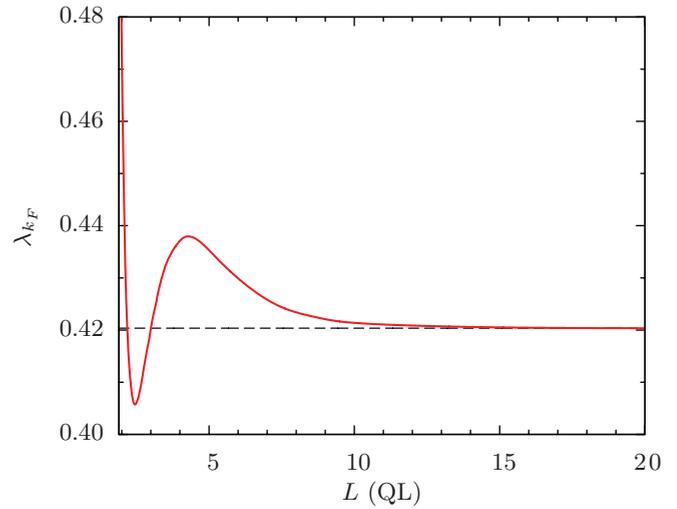


FIG. 6. (Color online) Width (L) dependence of the effective electron-phonon coupling constant at the Fermi level λ_{k_F} for $\mu = 0.2$ eV. The dashed horizontal line indicates one-half of the effective coupling constant in the semi-infinite geometry with otherwise identical parameters (Ref. 11).

For low temperatures and $k = k_F$, the decay rate is dominated by the $n = 1$ phonon mode with $q \rightarrow 0$. After some algebra, we find that this implies a T^2 law,

$$\Gamma_{k_F}(T \ll T_{\text{BG}}) = \frac{4\pi(c_t/c_l)^4 (k_F c_R \alpha)^2}{\rho_M |v_{k_F}| c_s^3} \frac{1}{L} \left(\frac{T}{T_{\text{BG}}} \right)^2. \quad (43)$$

Again, when $T \gtrsim T_c$, the T^2 law (which scales $\propto 1/L$) competes with the L -independent T^3 law found in Ref. 11; see Sec. III. Finally, when $k \neq k_F$ and $T \ll T_{\text{BG}}$, the quasiparticle decay rate saturates at the finite value

$$\Gamma_{k \neq k_F} = \alpha^2 \sum_n \int_0^\infty q dq \Theta(|\epsilon_{k+}| - \Omega_{q, n}) F_{k, n, +}^{(\nu)}(q) \quad (44)$$

with $\nu = \text{sgn}(k_F - k)$.

Figure 6 shows that the $L \rightarrow \infty$ limit of the decay rate always tends to $\Gamma_\infty(T)/2$, where Γ_∞ is the corresponding decay rate for the semi-infinite geometry.¹¹ This discrepancy with the naive expectation $\Gamma(L \rightarrow \infty) = \Gamma_\infty$ has the same origin as the anomalous factor 1/2 appearing in the large- L behavior of the resistivity discussed in Sec. III.

V. CONCLUSIONS

In this paper we have studied the effects of long-wavelength acoustic phonons on the topologically protected surface fermions in topological insulator films. Our model employs the established low-energy electronic Hamiltonian and an isotropic elastic continuum approach for the phonons, with the deformation coupling providing the dominant interaction mechanism. The full crossover from thin to thick films has been studied, taking into account transverse quantization effects. The electron-phonon coupling turns out to be surprisingly strong, in accordance with recent ARPES results.²⁴

Using a quasiclassical approach, we have computed the temperature-dependent resistivity of the film due to phonon backscattering, and found a linear T dependence above the

Bloch-Grüneisen temperature. In this temperature regime, the phonon-induced resistivity can overcome the disorder-induced (T -independent) contribution and should be observable with present samples. Similarly, the linear T dependence of the quasiparticle decay rate found here is observable²⁴ in ARPES experiments. The low-temperature behaviors of the resistivity and of the quasiparticle decay rate are probably more difficult to observe.

An interesting extension of our work would be to include the effects of a magnetic field. Magnetotransport measurements in thin films were recently performed⁴⁶ and found clear evidence for Landau level formation associated with the massless Dirac fermions forming on both surfaces. The observed broadening of the Landau levels was assigned to disorder and/or interaction

effects, but at elevated temperatures, our analysis indicates that electron-phonon interactions may be relevant as well.

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APPENDIX A: NORMALIZATION CONSTANTS

Here we provide the normalization constants $\mathcal{N}_{S,A}$ appearing in Eqs. (22) and (23) in Sec. II B. Specifically, we get these constants after some algebra from

$$\begin{aligned}\mathcal{N}_S^{-2} &= \frac{\Omega_\Lambda^2}{2c_t^2} \left[\frac{c_t^2}{c_l^2} \frac{4q^2 L}{\cos^2(k_l L/2)} \left(1 + \frac{\sin(k_l L)}{k_l L} \right) + \frac{(q^2 - k_t^2)^2 L}{k_t^2 \cos^2(k_l L/2)} \left(1 - \frac{\sin(k_l L)}{k_l L} \right) - 8(q^2 - k_t^2) \frac{\tan(k_l L/2)}{k_t} \right], \\ \mathcal{N}_A^{-2} &= \frac{\Omega_\Lambda^2}{2c_t^2} \left[\frac{c_t^2}{c_l^2} \frac{4q^2 L}{\sin^2(k_l L/2)} \left(1 - \frac{\sin(k_l L)}{k_l L} \right) + \frac{(q^2 - k_t^2)^2 L}{k_t^2 \sin^2(k_l L/2)} \left(1 + \frac{\sin(k_l L)}{k_l L} \right) + 8(q^2 - k_t^2) \frac{\cot(k_l L/2)}{k_t} \right].\end{aligned}$$

APPENDIX B: LINEARIZED BOLTZMANN EQUATION

The linearized Boltzmann equation has been derived for closely related problems before,^{30,35,36,44} and we here follow those works and briefly sketch the derivation of Eq. (28). In the quasiclassical approximation, the quasiparticle distribution function $f(\mathbf{r}, \mathbf{k}, t)$ (for simplicity, we here omit the ν and s indices) obeys the well-known Boltzmann equation. In the absence of a force $\mathbf{F} = -e\mathbf{E}$ due to the external electric field, f reduces to a Fermi function, $f = n_F(\epsilon_{\mathbf{k}})$, and the collision integral vanishes. In the presence of the force, f is expanded in terms of Legendre polynomials $P_n(\cos\alpha)$, where α is the angle between \mathbf{k} and \mathbf{F} . Keeping only terms linear in \mathbf{F} , we have $f(\mathbf{k}) = n_F(\epsilon_{\mathbf{k}}) + \cos(\alpha) f_1(\epsilon_{\mathbf{k}})$. Using detailed balance, for given transition matrix elements $W_{\mathbf{k} \rightarrow \mathbf{k}'}$, we obtain the linearized Boltzmann equation (LBE),

$$\begin{aligned}\mathbf{F} \cdot \mathbf{v}_{\mathbf{k}} \partial_\epsilon n_F(\epsilon_{\mathbf{k}}) &= \sum_{\mathbf{k}'} W_{\mathbf{k} \rightarrow \mathbf{k}'} \left[\frac{n_F(\epsilon_{\mathbf{k}})}{n_F(\epsilon_{\mathbf{k}'})} \cos(\alpha') f_1(\epsilon_{\mathbf{k}'}) \right. \\ &\quad \left. - \frac{1 - n_F(\epsilon_{\mathbf{k}'})}{1 - n_F(\epsilon_{\mathbf{k}})} \cos(\alpha) f_1(\epsilon_{\mathbf{k}}) \right].\end{aligned}$$

Using the ansatz $f_1(\epsilon_{\mathbf{k}}) = -\tau(\epsilon_{\mathbf{k}}) v_{\mathbf{k}} F \partial_\epsilon n_F(\epsilon_{\mathbf{k}})$, after some algebra the LBE leads to the linear integral equation

$$\frac{1}{\tau(\epsilon_{\mathbf{k}})} = \sum_{\mathbf{k}'} W_{\mathbf{k} \rightarrow \mathbf{k}'} \left[1 - \frac{v_{\mathbf{k}'} \tau(\epsilon_{\mathbf{k}'})}{v_{\mathbf{k}} \tau(\epsilon_{\mathbf{k}})} \cos(\vartheta) \right] \frac{1 - n_F(\epsilon_{\mathbf{k}'})}{1 - n_F(\epsilon_{\mathbf{k}})},$$

where ϑ is the angle between \mathbf{k} and \mathbf{k}' . The solution for $\tau(\epsilon_{\mathbf{k}})$ determines the electron momentum relaxation time. When the scattering of quasiparticles from long-wavelength acoustic phonons is quasielastic, $\Omega_{q,n} \ll |\mu|$, we can set $\tau(\epsilon_{\mathbf{k}'}) = \tau(\epsilon_{\mathbf{k}})$ for the right-hand side of the above integral equation; this is equivalent to the ‘‘test particle approximation.’’³⁰ The current density $\mathbf{j} = -(e/A) \sum_{\mathbf{k}} \mathbf{v}_{\mathbf{k}} f(\mathbf{k})$ points parallel to the electric field direction and has the magnitude

$$j = \frac{e^2 E}{\mathcal{A}} \sum_{\mathbf{k}} v_{\mathbf{k}}^2 \cos^2(\alpha) \tau(\epsilon_{\mathbf{k}}) [-\partial_\epsilon n_F(\epsilon_{\mathbf{k}})].$$

Using $v_{\mathbf{k}} = v_k$ and performing the angular integration, we arrive at the phonon contribution to the resistivity quoted in Eqs. (28) and (29).

APPENDIX C: TRANSPORT ELIASHBERG FUNCTION

We here give the analytical result for the full transport Eliashberg function \mathcal{F} defined in Eq. (32). Some straightforward yet tedious algebra allows one to perform the θ integration. We find the (lengthy) result

$$\begin{aligned}\mathcal{F}_{k,n,s}^{(v)}(q) &= \frac{2|A_{k,q,n,s}^{(v)}|}{\pi(A_2 W)^2} \frac{\Theta(Q_{k,q,n,s}^{(v)} + k - q) \Theta(q - |Q_{k,q,n,s}^{(v)} - k|)}{[(q^2 - (Q_{k,q,n,s}^{(v)} - k)^2)((Q_{k,q,n,s}^{(v)} + k)^2 - q^2)]^{1/2}} \Theta(|A_{k,q,n,s}^{(v)}| - \Delta/2) \\ &\quad \times \sum_{s'} \Theta(s' A_{k,q,n,s}^{(v)}) \left[1 - \left(1 - \frac{\nu \Omega_{q,n}}{A_{k,q,n,s}^{(v)}} \right) \frac{(Q_{k,q,n,s}^{(v)})^2 + k^2 - q^2}{2k^2} \right] \Big|_{\theta_0}^{s',s},\end{aligned}$$

where we use the notations

$$A_{k,q,n,s}^{(v)} = s\sqrt{(\Delta/2)^2 + (A_2 W k)^2} + v\Omega_{q,n}, \quad Q_{k,q,n,s}^{(v)} = \frac{\sqrt{(A_{k,q,n,s}^{(v)})^2 - (\Delta/2)^2}}{A_2 W},$$

and the polar angle $\theta = \theta_0 \in [0, \pi]$ follows from

$$\sqrt{k^2 + q^2 + 2kq \cos \theta_0} = Q_{k,q,n,s}^{(v)},$$

fixing the polar angle between \mathbf{k} and \mathbf{q} in the matrix element M .

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$\kappa_l^2 \kappa_l \rightarrow \kappa_l \kappa_l^2$. This implies $C \simeq 0.78$ instead of the quoted value $C \simeq 1.20$ in Ref. 11. Fortunately, this mistake has no consequences for the conclusions reached in Ref. 11.

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